

# Mass hierarchies in string theory and holography

I. Antoniadis

CERN

SUPERFIELDS - Sixth Joint Seminar

Padova 18 March 2011

- ① Motivations and mass hierarchy
- ② Strings, branes and extra dimensions
- ③ Gravity scale and number of species
- ④ Low string coupling and little strings at a TeV

I.A.-Dimopoulos-Giveon '01 + with Arvanitaki '11

- ⑤ Main accelerator signatures

# BSM physics: driven by mass hierarchy problem

Higgs mass: very sensitive to high energy physics       $m_H \sim \text{UV cutoff } \Lambda$

why gravity is so weak compared to the other interactions?       $\Lambda = M_P$

Possible answer (alternative to supersymmetry): Low UV cutoff       $\Lambda \sim \text{TeV}$

- low scale gravity  $\Rightarrow$  large extra dimensions, warped dimensions
- low string scale  $\Rightarrow$  low scale gravity, ultra weak string coupling

Experimentally testable framework:

- spectacular model independent predictions
  - radical change of high energy physics at the TeV scale
- explicit model building is not necessary at this moment

# Framework of type I string theory $\Rightarrow$ D-brane world

I.A.-Arkani-Hamed-Dimopoulos-Dvali '98

- gravity: closed strings propagating in 10 dims
- gauge interactions: open strings with their ends attached on D-branes

Dimensions of finite size:  $n$  transverse       $6 - n$  parallel

calculability  $\Rightarrow R_{\parallel} \simeq l_{\text{string}}$ ;  $R_{\perp}$  arbitrary

$$M_P^2 \simeq \frac{1}{g_s^2} M_s^{2+n} R_{\perp}^n \quad g_s = \alpha : \text{weak string coupling}$$

Planck mass in  $4 + n$  dims:  $M_*^{2+n}$

$$M_s \sim 1 \text{ TeV} \Rightarrow R_{\perp}^n = 10^{32} l_s^n \quad \text{small } M_s/M_P : \text{extra-large } R_{\perp}$$

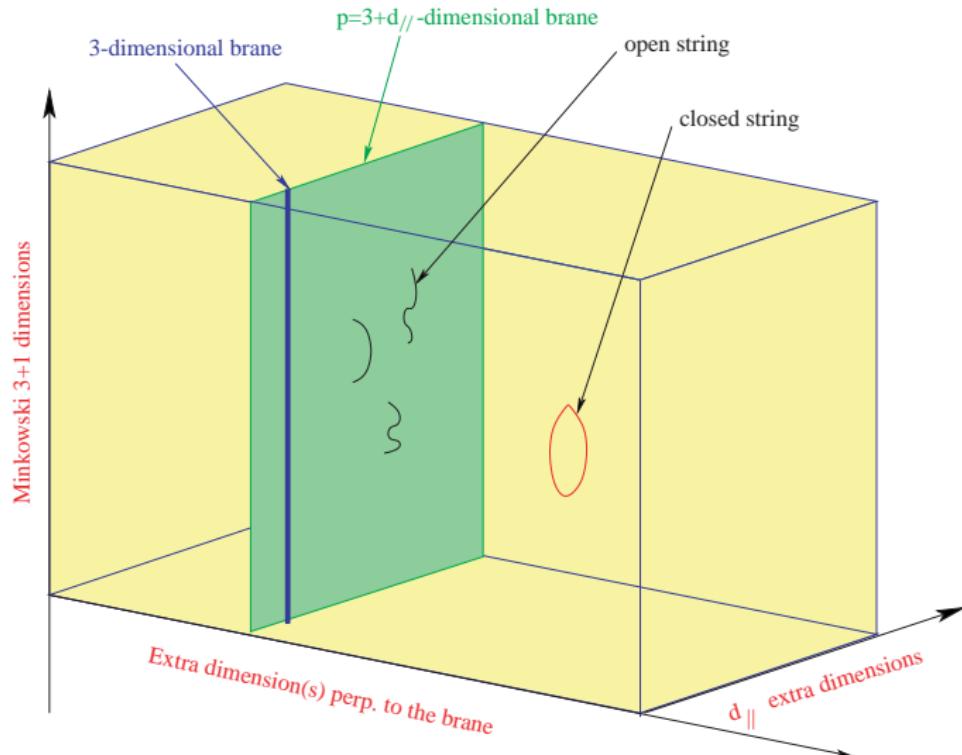
$$R_{\perp} \sim .1 - 10^{-13} \text{ mm for } n = 2 - 6 \quad [5]$$

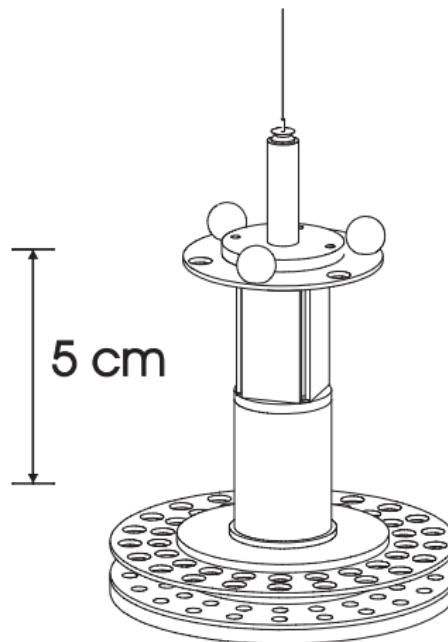
distances  $< R_{\perp}$ : gravity  $(4+n)$ -dim  $\rightarrow$  strong at  $10^{-16}$  cm [6]

# Braneworld

2 types of compact extra dimensions:

- parallel ( $d_{\parallel}$ ):  $\lesssim 10^{-16}$  cm (TeV) I.A. '90
- transverse ( $\perp$ ):  $\lesssim 0.1$  mm (meV)

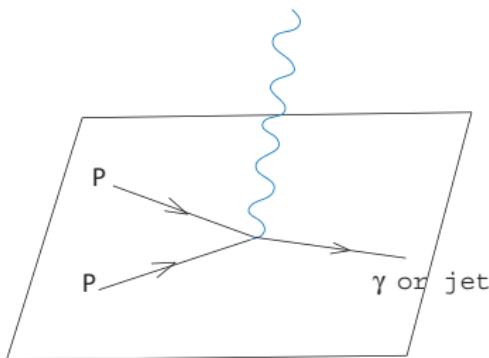




$R_{\perp} \lesssim 45 \mu\text{m}$  at 95% CL

- dark-energy length scale  $\approx 85 \mu\text{m}$  [3]

# Gravitational radiation in the bulk $\Rightarrow$ missing energy



Collider bounds on $R_{\perp}$ in mm			
	$n = 2$	$n = 4$	$n = 6$
LEP 2	$4.8 \times 10^{-1}$	$1.9 \times 10^{-8}$	$6.8 \times 10^{-11}$
Tevatron	$5.5 \times 10^{-1}$	$1.4 \times 10^{-8}$	$4.1 \times 10^{-11}$
LHC	$4.5 \times 10^{-3}$	$5.6 \times 10^{-10}$	$2.7 \times 10^{-12}$
NLC	$1.2 \times 10^{-2}$	$1.2 \times 10^{-9}$	$6.5 \times 10^{-12}$

# Black hole production

String-size black hole energy threshold :  $M_{\text{BH}} \simeq M_s/g_s^2$

Horowitz-Polchinski '96, Meade-Randall '07

- string size black hole:  $r_H \sim l_s = M_s^{-1}$
- black hole mass:  $M_{\text{BH}} \sim r_H^{d-3}/G_N \quad G_N \sim l_s^{d-2} g_s^2$

weakly coupled theory  $\Rightarrow$  strong gravity effects occur much above  $M_s, M_*$   
 $g_s \sim 0.1$  (gauge coupling)  $\Rightarrow M_{\text{BH}} \sim 100M_s$

Comparison with Regge excitations :  $M_n = M_s \sqrt{n} \Rightarrow$   
production of  $n \sim 1/g_s^4 \sim 10^4$  string states before reach  $M_{\text{BH}}$  [3]

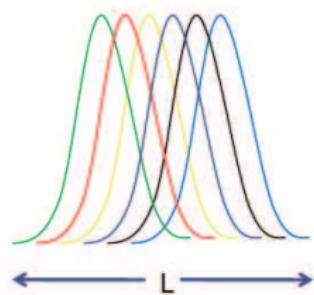
# More general framework: large number of species

$N$  particle species  $\Rightarrow$  lower quantum gravity scale :  $M_*^2 = M_p^2/N$

Dvali '07, Dvali, Redi, Brustein, Veneziano, Gomez, Lüst '07-'10

derivation from: black hole evaporation or quantum information storage

Pixel of size  $L$  containing  $N$  species storing information:



localization energy  $E \gtrsim N/L \rightarrow$

Schwarzschild radius  $R_s = N/(LM_p^2)$

no collapse to a black hole :  $L \gtrsim R_s \Rightarrow L \gtrsim \sqrt{N}/M_p = 1/M_*$

$M_* \simeq 1 \text{ TeV} \Rightarrow N \sim 10^{32} \text{ particle species !}$

## 2 ways to realize $N = 10^{32}$ lowering the string scale

- ① Large volume compactifications      SM on D-branes

$N = R_{\perp}^n I_s^n$  : number of KK modes up to energies of order  $M_* \simeq M_s$

- ②  $N \sim$  effective number of string modes contributing to the BH bound

Dvali-Lüst '09, Dvali-Gomez '10

$$N_s = \frac{1}{g_s^2} \text{ with } g_s \simeq 10^{-16} \quad \text{SM on NS5-branes}$$

I.A.-Pioline '99, I.A.-Dimopoulos-Giveon '01

in this case gravity does NOT become strong at  $M_s$

Both ways are compatible with the general string relation:

$$M_p^2 = \frac{1}{g_s^2} V_6 M_s^8 \quad V_6 : \text{internal 6d compactification volume}$$

puzzle as  $g_s \rightarrow 0$ :

- $M_*$  remains finite  $\Rightarrow$  Quantum Gravity effects in a free theory

$$M_*^2 = M_p^2/N \simeq M_s^2/(g_s^2 N) \sim M_s^2 \quad \text{since } N \simeq 1/g_s^2$$

- forward 4pt amplitude does not decouple

e.g. 
$$\sum_{\text{string states } X} |q\bar{q} \rightarrow X|^2 \sim g_s^2 \times N \sim \mathcal{O}(1) \text{ at } M_*$$

solution: log corrections  $N(M_*) \sim 1/(g_s \ln g_s)^2$   $M_* \sim M_s |\ln g_s|$

string density of states:  $N(M_*) \sim \left(\frac{M_*}{M_s}\right)^{-d} e^{\beta M_*/M_s}$

$d = 4$ ;  $\beta = 2\sqrt{2}\pi$  for closed superstrings

$$M_*^2 = \frac{M_s^2}{g_s^2 N} v_6 \leftarrow \text{6d volume in string units}$$

$$\Rightarrow \frac{1}{g_s \sqrt{N}} = \frac{1}{\beta} \ln N - \frac{d}{\beta} \ln \ln N + \dots$$

# What is LST ? Decouple gravity from NS5-branes

Analogy from D3-branes : decouple gravity  $\Rightarrow M_s \rightarrow \infty$ ,  $g_s$  fixed  
 $\rightarrow$  (conformal) Field Theory (CFT)

simplest case: 4d  $\mathcal{N} = 4$  super Yang Mills  $SU(N)$

parameters: number of branes  $N$ , gauge coupling  $g_{YM}$

NS-5 branes:  $M_s$  finite,  $g_s \rightarrow 0 \rightarrow$  (little) String Theory without gravity

simplest case: 6d LST (chiral IIA or non-chiral IIB)

massless sector: 6d  $SU(N)$  of vectors (IIA) or tensors (IIB)

at a non-trivial fixed point

parameters: number of branes  $N$ , string scale  $M_s$

# How to study LST ? Using gauge/gravity duality

Gravity background : near horizon geometry (holography) Maldacena '98

Analogy from D3-branes :  $AdS_5 \times S^5$

parameters:  $AdS$  radius  $r_{AdS} M_s$ ,  $g_s \leftrightarrow N, g_{YM}$

supergravity validity:  $r_{AdS} M_s \gg 1, g_s \ll 1 \Rightarrow$  large  $N, g_{YM}^2 N$

→ model independent part :  $AdS_5$

NS-5 branes :  $(\mathcal{M}_6 \otimes R_+) \times SU(2) \equiv S^3$



linear dilaton background in 7d flat string-frame metric  $\Phi = -\alpha|y|$

Aharony-Berkooz-Kutasov-Seiberg '98

parameters:  $M_s, \alpha$  (or  $S^3$  radius)  $\leftrightarrow N$

sugra validity: small  $\alpha \Rightarrow$  large  $N$

compactify to  $d = 4$  ( $\mathcal{M}_6 \rightarrow \mathcal{M}_4$ )  $\Rightarrow g_{YM} \sim 2d$  volume

→ model independent part : linear dilaton

# Put gravity back but weakly coupled

“cut” the space of the extra dimension  $\Rightarrow$  gravity on the brane

Analogy from D-branes  $\rightarrow$  2 possibilities:

- flat space  $\Rightarrow$  large extra dimensions AADD '98
- curved space from gravity back reaction  $\Rightarrow$  slice of  $AdS_5$  RS, H. Verlinde '99

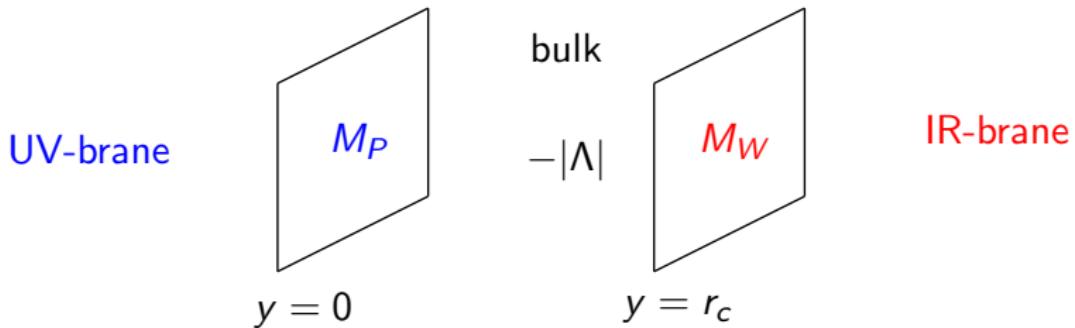
NS-5 branes : linear dilaton on an interval  $y \in [0, r_c]$

$$S_{bulk} = \int d^4x dy \sqrt{-g} e^{-\Phi} (M_5^3 R + M_5^3 (\nabla \Phi)^2 - \Lambda)$$

$$S_{vis(hid)} = \int d^4x \sqrt{-g} e^{-\Phi} (L_{SM(hid)} - T_{vis(hid)})$$

Tuning conditions:  $T_{vis} = -T_{hid} \leftrightarrow \Lambda < 0$  [16]

RS models :  $ds^2 = e^{-2k|y|} \eta_{\mu\nu} dx^\mu dx^\nu + dy^2$      $k^2 \sim \Lambda/M_5^3$



- fine-tuned tensions:  $T = -T' = 24M^3 k$
- exponential hierarchy:  $M_W = M_P e^{-2kr_c}$      $M_P^2 \sim M_5^3/k$   
 $M_5 \sim M_{GUT}$
- 4d gravity localized on the UV-brane, but KK gravitons on the IR

- main prediction: spin-2 resonances at the TeV scale

$$m_n = c_n k e^{-2kr_c} \sim \text{TeV} \quad c_n \simeq (n + 1/4) \text{ for large } n$$

⇒ spin-2 TeV resonances in di-lepton or di-jet channels

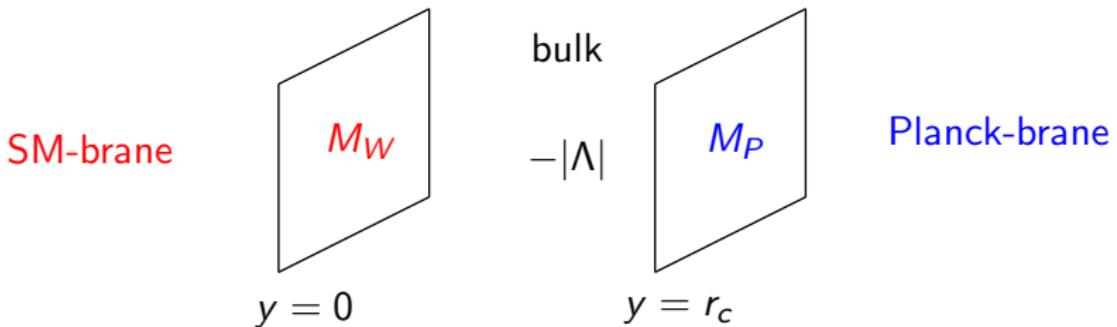
weakly coupled for  $k < M_5$

- viable models: SM gauge bosons in the bulk, Higgs on the IR-brane
- AdS/CFT duals to strongly coupled 4d field theories

composite Higgs models, technicolor-type     $g_{YM} = M_5/k > 1$  [13]

$$\text{LST} : \Phi = -\alpha|y| ; ds^2 = e^{\frac{2}{3}\alpha|y|} (\eta_{\mu\nu} dx^\mu dx^\nu + dy^2) \leftarrow \text{Einstein frame}$$

$z \sim e^{\alpha y/3} \Rightarrow$  polynomial warp factor + log varying dilaton

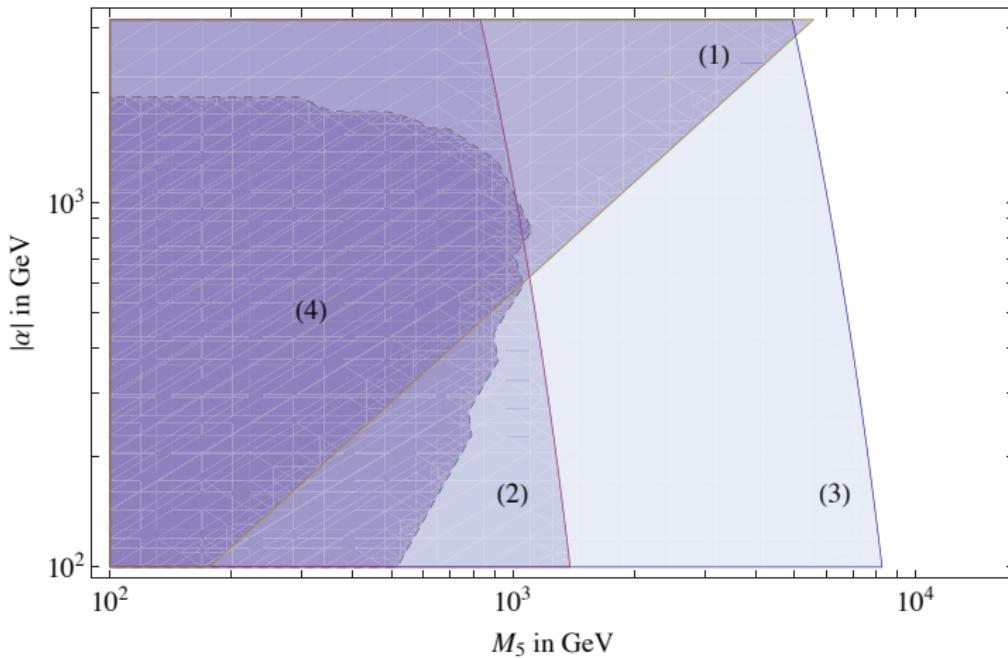


- exponential hierarchy:  $g_s^2 = e^{-\alpha|y|} \quad M_P^2 \sim \frac{M_5^3}{\alpha} e^{\alpha r_c}$
- 4d graviton flat, KK gravitons localized near SM
- SM particles cannot be in the bulk  
bulk gauge bosons: exp suppressed couplings

## LST KK graviton phenomenology

- KK spectrum :  $m_n^2 = \left(\frac{n\pi}{r_c}\right)^2 + \frac{\alpha^2}{4}$  ;  $n = 1, 2, \dots$   
⇒ mass gap + dense KK modes       $\alpha \sim 1 \text{ TeV}$        $r_c^{-1} \sim 30 \text{ GeV}$
- couplings :  $\frac{1}{\Lambda_n} \sim \frac{1}{(\alpha r_c) M_5}$   
⇒ extra suppression by a factor  $(\alpha r_c) \simeq 30$
- width :  $1/(\alpha r_c)^2$  suppression  $\sim 1 \text{ GeV}$   
⇒ narrow resonant peaks in di-lepton or di-jet channels
- extrapolates between RS and flat extra dims ( $n = 1$ )  
 $\alpha \gtrsim (0.1 \text{ mm})^{-1} \sim 10^{-2} \text{ eV}$  from microgravity experiments  
⇒ distinct experimental signals

# Bounds on the LST parameter space



exclusion by (1) perturbativity (2) Tevatron with  $5.4 \text{ fb}^{-1}$  data

(3) LHC 14 TeV with  $10 \text{ fb}^{-1}$  (4) diphoton at Tevatron  $5.4 \text{ fb}^{-1}$

# Conclusions

Mass hierarchy  $\Rightarrow$  testing strings at the TeV ?

- Well motivated theoretical framework
  - with many testable experimental predictions
  - new resonances, missing energy
- Several realizations with different signatures
  - flat large extra dimensions, exp warped metrics,
  - tiny string coupling and linear dilaton background
- Stimulus for micro-gravity experiments and accelerator searches

But: - unification has to be dropped

- physics is radically changed above string scale